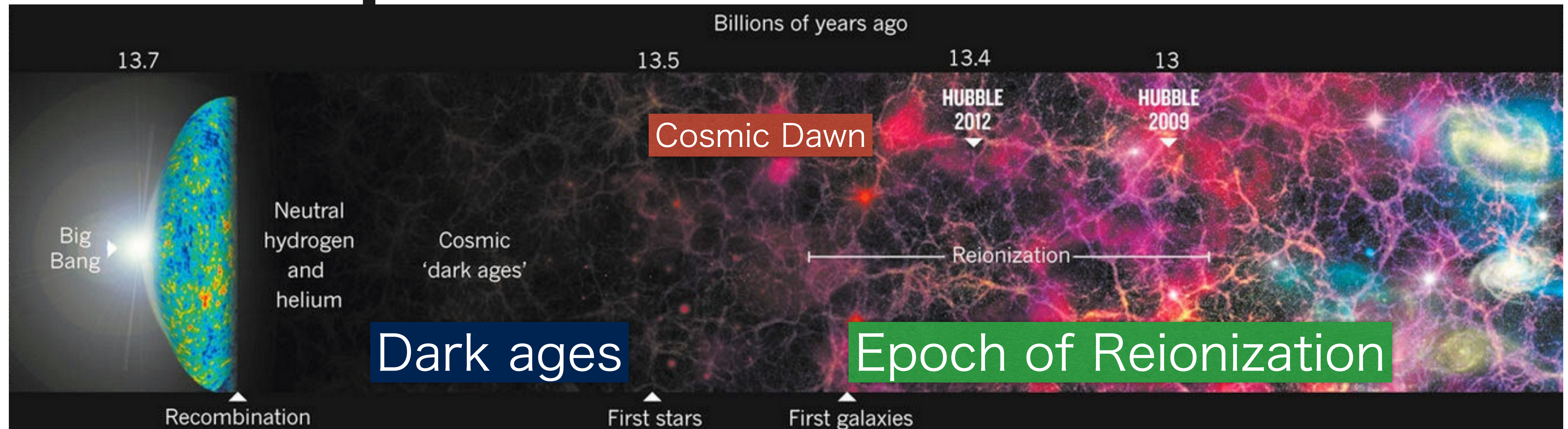


# Foreground Removal with Gaussian Process Regression

MWA Project Meeting -2023

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# What is the Epoch of Reionization



<https://astrobites.org/wp-content/uploads/2015/05/cover.png>

## Dark ages

→ After the cosmic recombination, and there have been no luminous objects

## Cosmic Dawn

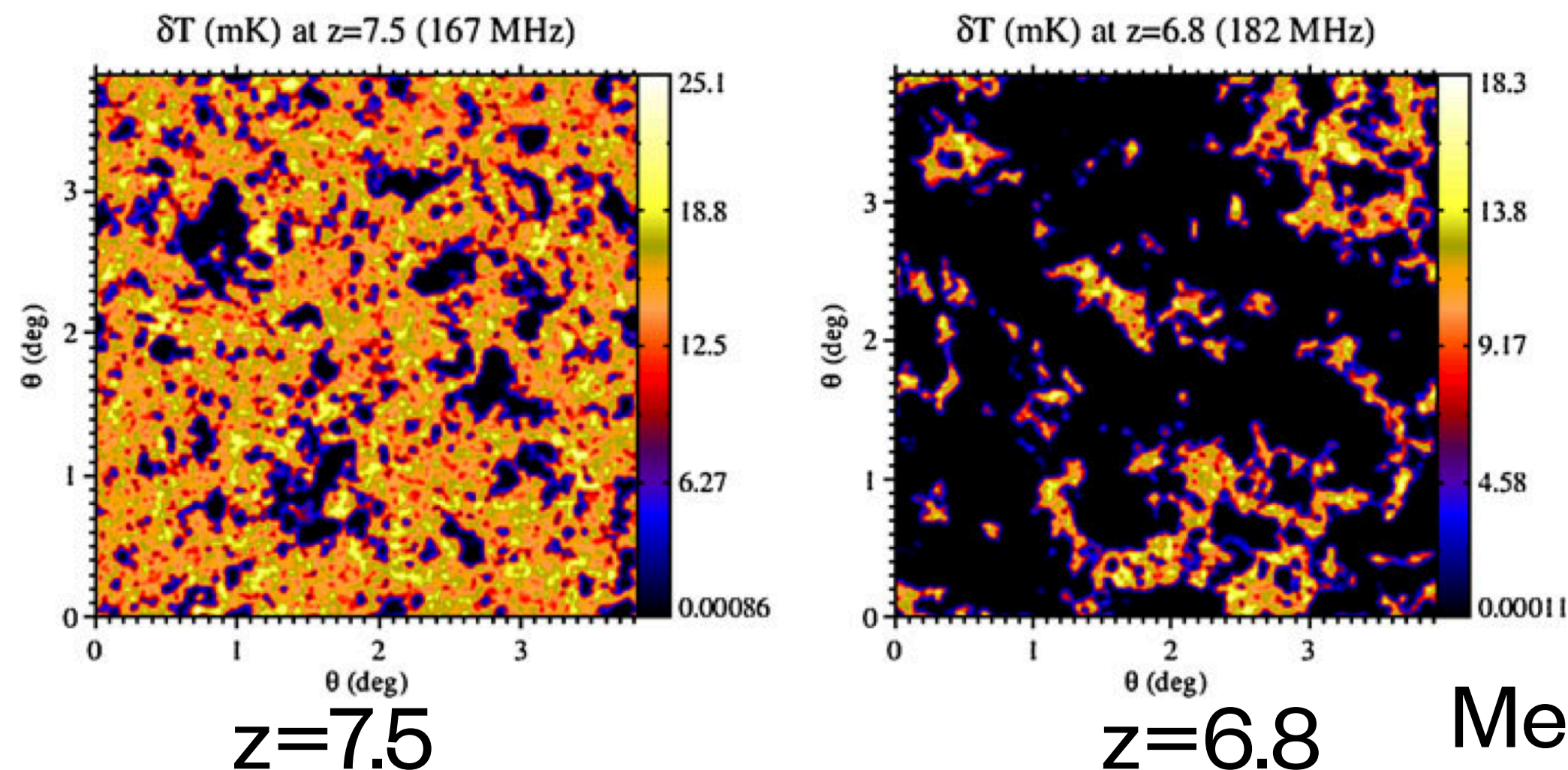
→ First luminous objects are formed.

## Epoch of Reionization

→ ionizing photons from galaxies ionized the neutral hydrogen gas distributed in the Universe.

# 21 cm line

- The 21 cm line emission is due to the spin flip of HI  
→ We can observe IGM at the EoR via 21 cm line
- HI distribution at different redshifts can be observed by different frequencies  
→ We can follow the evolution of IGM



IGM simulation

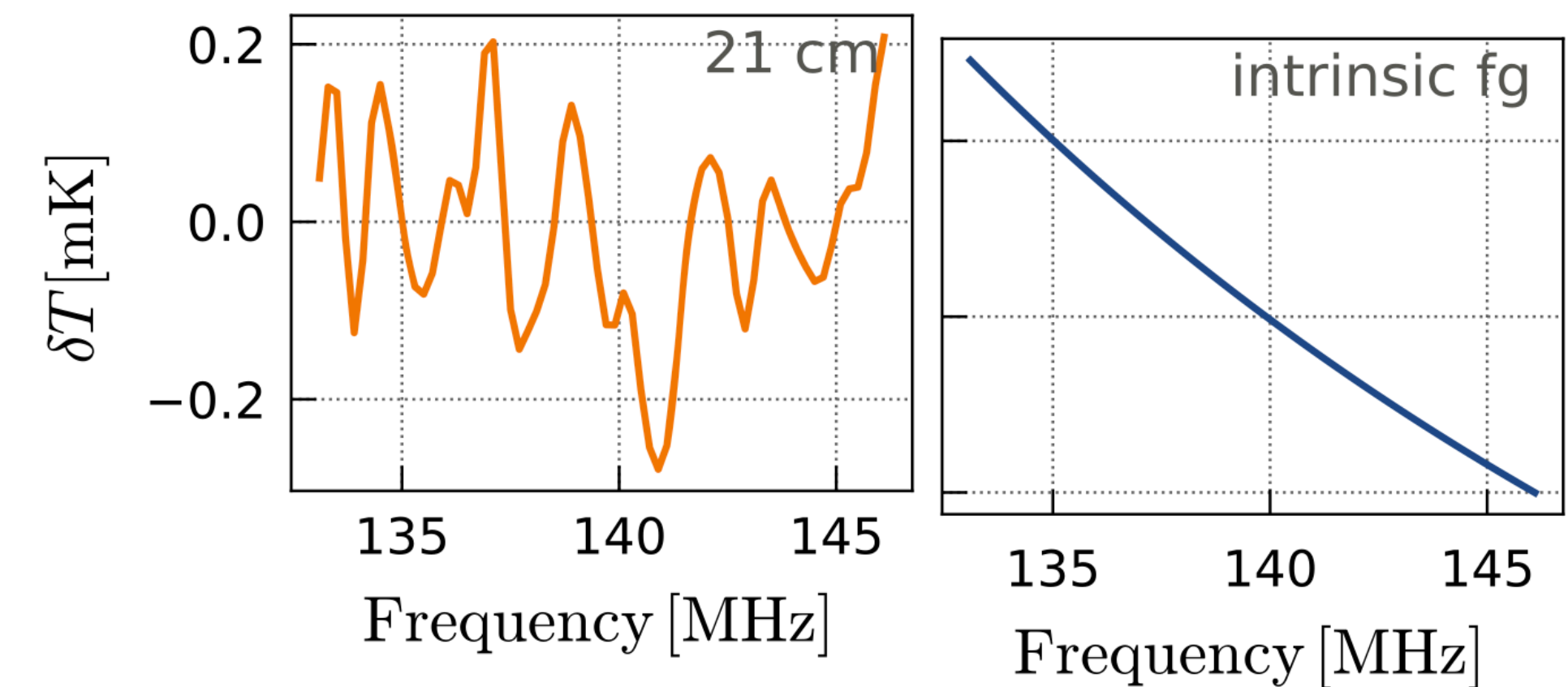
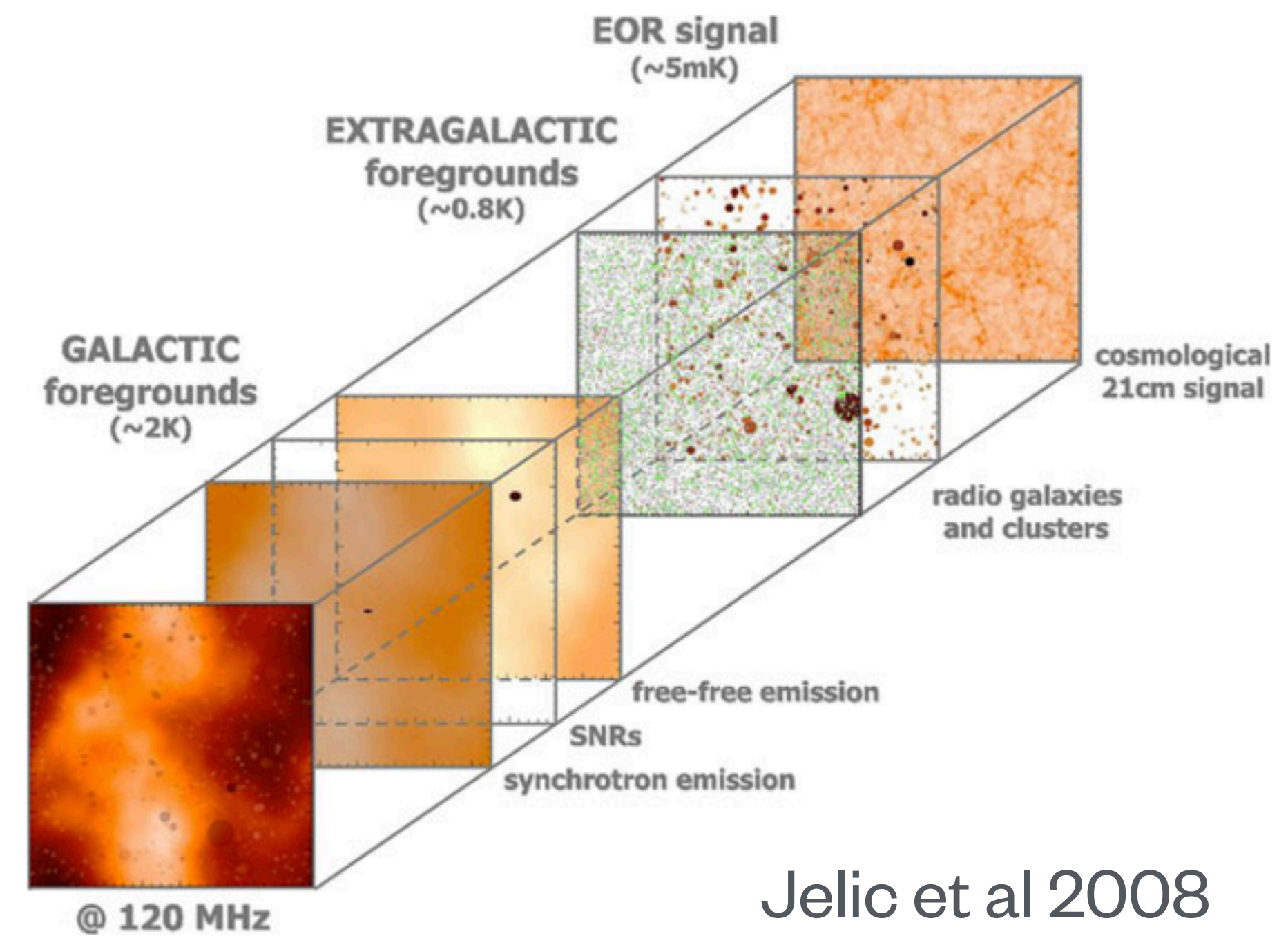
orange: HI emission (not ionized)

black: no HI emission (ionized)

Mellema et al 2013

# Foregrounds

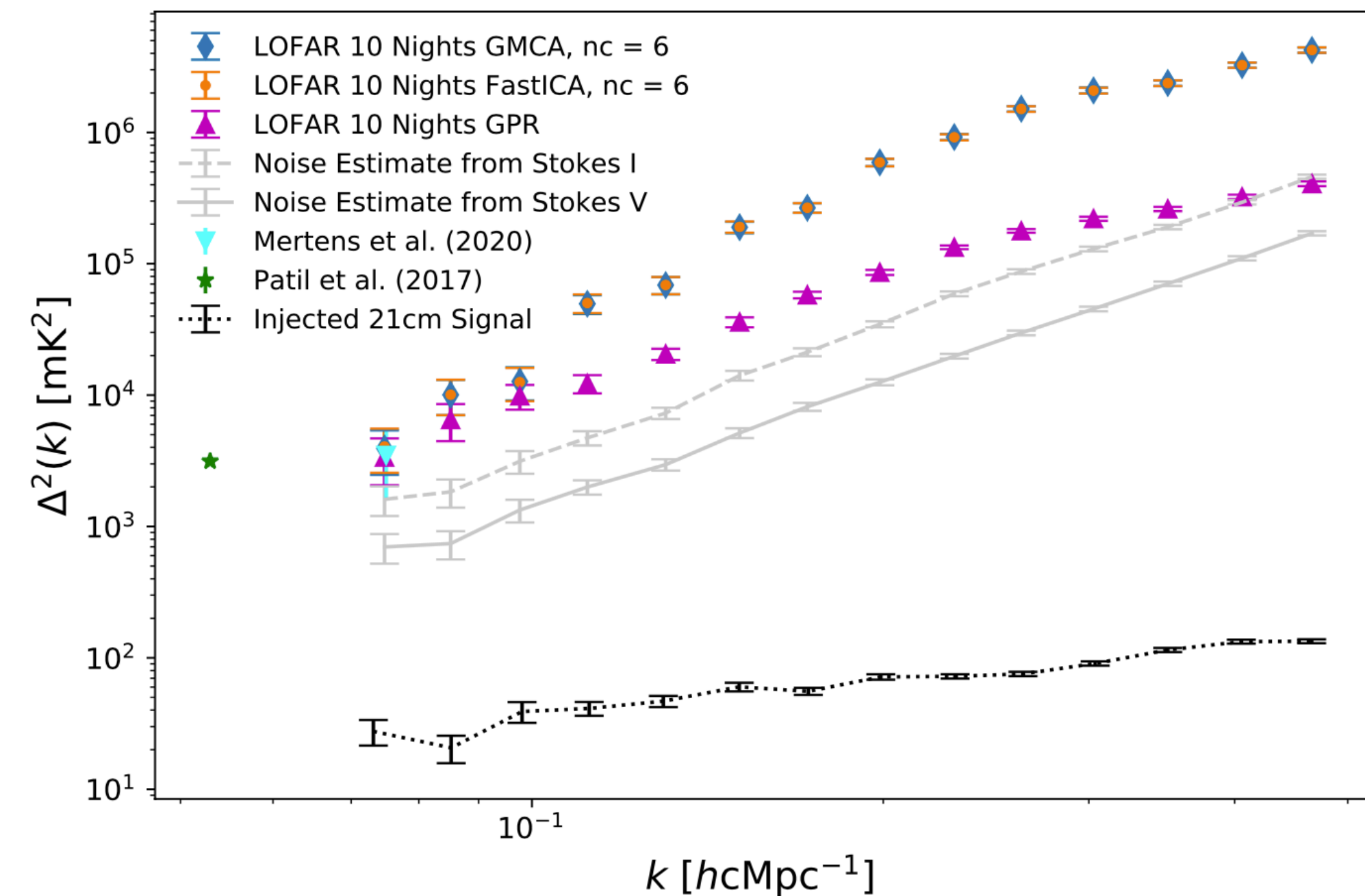
- Observed signal
  - Foreground(FG) + 21cm line + Noise
- FG is brighter than EoR signal
  - ( $\sim 10^3$  in order)
- Avoidance or Removal of FG is important
  - How to remove foreground?
  - Use difference between FG and EoR signal
    - Emission strength(FG  $\gg$  EoR signal)
    - Spectral behavior



Mertens et al 2018

# Foreground Removal techniques

- There are various foreground removal techniques
  - Generalized Morphological Component Analysis (GMCA)
  - FAST ICA
  - Principal Component Analysis (PCA)
  - Gaussian Process Regression (GPR)
- Hothi et al 2020 reports that GPR has better performance than FastICA, GMCA



# My research

- GPR uses covariance called kernel to represent data
  - Best kernels set maybe different for each telescope
  - Compare some of the kernels combination by Bayesian evidence
- Imaging might be affect to FG removal
  - Visibility based FG removal(LOFAR analysis is image based)

I applied Gaussian Process Regression based FG removal to MWA observational data

# Gaussian Process Regression(GPR)

- Our data consists of the foreground ( $\mathbf{f}_{fg}$ ), EoR signal ( $\mathbf{f}_{21}$ ) and noise ( $\mathbf{n}$ )

$$\mathbf{d} = \mathbf{f}_{fg} + \mathbf{f}_{21} + \mathbf{n}$$

- Assuming each component to be statistically uncorrelated, the covariance of the data  $\mathbf{K}$  is given by

$$\mathbf{K} = \mathbf{K}_{fg} + \mathbf{K}_{21} + \mathbf{K}_{noise}$$

# Gaussian Process Regression(GPR)

- Gaussian Process(GP)→Multivariate Gaussian Distribution,  $N$

- If we assume random value  $f$  follows GP, we write

$$f \sim N(m, K)$$

where  $m$  :mean,  $K$ : covariance(kernel)



# Gaussian Process Regression(GPR)

- Assuming the data  $\mathbf{d}$  is Gaussian distributed, we can model its probability distribution as

$$\mathbf{d} \sim N(m(\nu), \mathbf{K}(\nu, \nu))$$

- We can write joint probability distribution of the GP at a series of other points in space  $\nu'$  as

$$\begin{bmatrix} \mathbf{d} \\ \mathbf{d}' \end{bmatrix} \sim N \left( \begin{bmatrix} m(\nu) \\ m(\nu') \end{bmatrix}, \begin{bmatrix} \mathbf{K}(\nu, \nu) & \mathbf{K}(\nu, \nu') \\ \mathbf{K}(\nu', \nu) & \mathbf{K}(\nu', \nu') \end{bmatrix} \right)$$

# Gaussian Process Regression(GPR)

- for the case of foreground removal, we want to estimate foreground model  $\mathbf{f}_{fg}$

$$\begin{bmatrix} \mathbf{d} \\ \mathbf{f}_{fg} \end{bmatrix} \sim \left( \begin{bmatrix} 0 \\ 0 \end{bmatrix}, \begin{bmatrix} \mathbf{K}_{fg} + \mathbf{K}_{21} + \mathbf{K}_n & \mathbf{K}_{fg} \\ & \mathbf{K}_{fg} \end{bmatrix} \right)$$

$$\mathbf{E}[\mathbf{f}_{fg}] = \mathbf{K}_{fg}[\mathbf{K}_{fg} + \mathbf{K}_{21} + \mathbf{K}_n]^{-1} \mathbf{d}$$

$$\text{Cov}(\mathbf{f}_{fg}) = \mathbf{K}_{fg} - \mathbf{K}_{fg}[\mathbf{K}_{fg} + \mathbf{K}_{21} + \mathbf{K}_n]^{-1} \mathbf{K}_{fg}$$

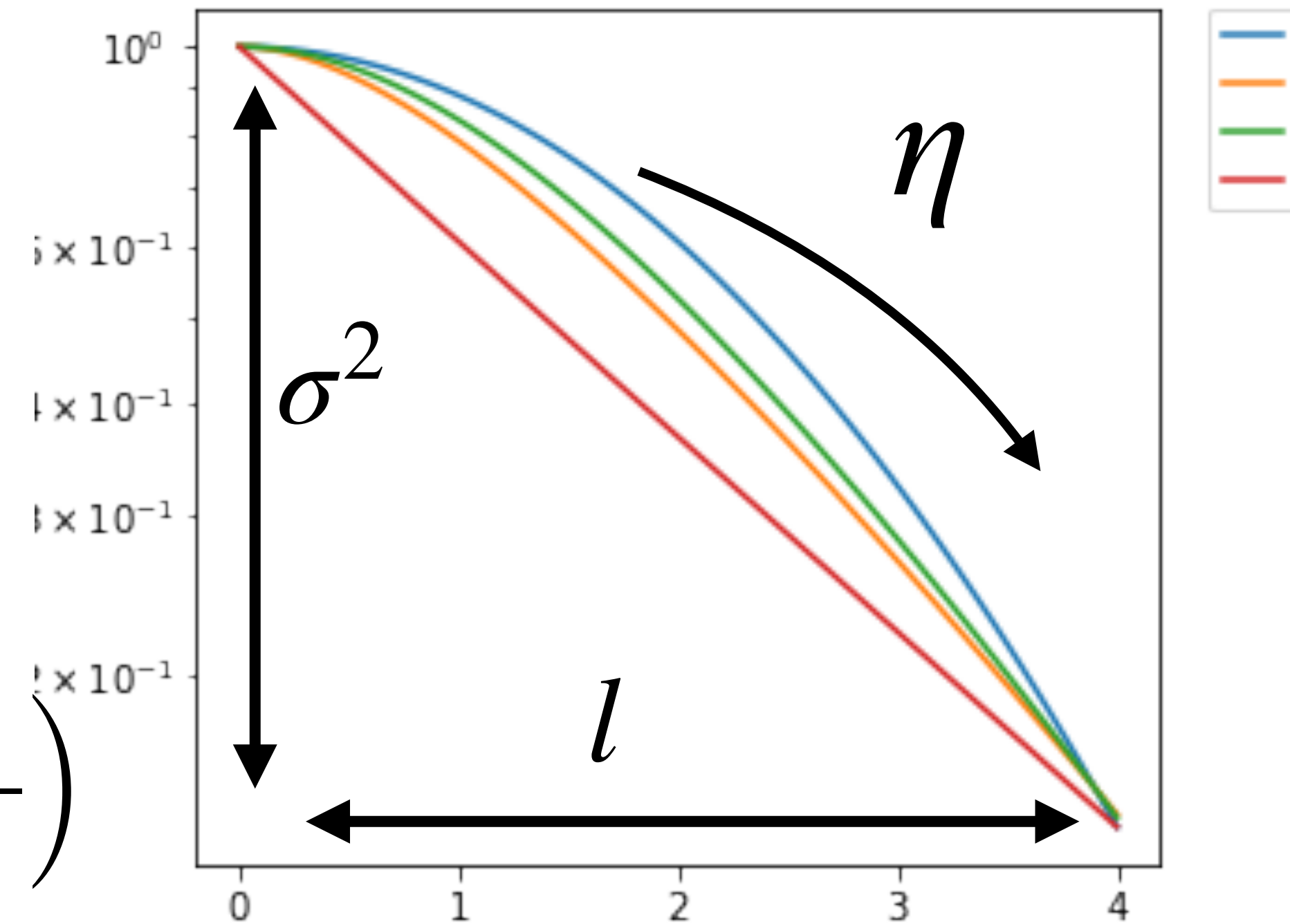
- To remove FG, subtract expected value of FG  $\mathbf{E}[\mathbf{f}_{fg}]$  from the data

$$\text{Residual} = \mathbf{d} - \mathbf{E}[\mathbf{f}_{fg}]$$

# Covariance (Kernel)

- Matern kernel is widely used kernels in GPR

$$K_{\text{Matern}}(\nu, \nu') = \sigma^2 \frac{2^{1-\eta}}{\Gamma(\eta)} \left( \sqrt{2\eta} \frac{|\nu - \nu'|}{l} \right)^\eta K_\eta \left( \sqrt{2\eta} \frac{|\nu - \nu'|}{l} \right)$$



$\Gamma$ :gamma function,  $K_\eta$ :modified Bessel function of the second kind,

$\sigma^2$ :Variance(amplitude of the signal)

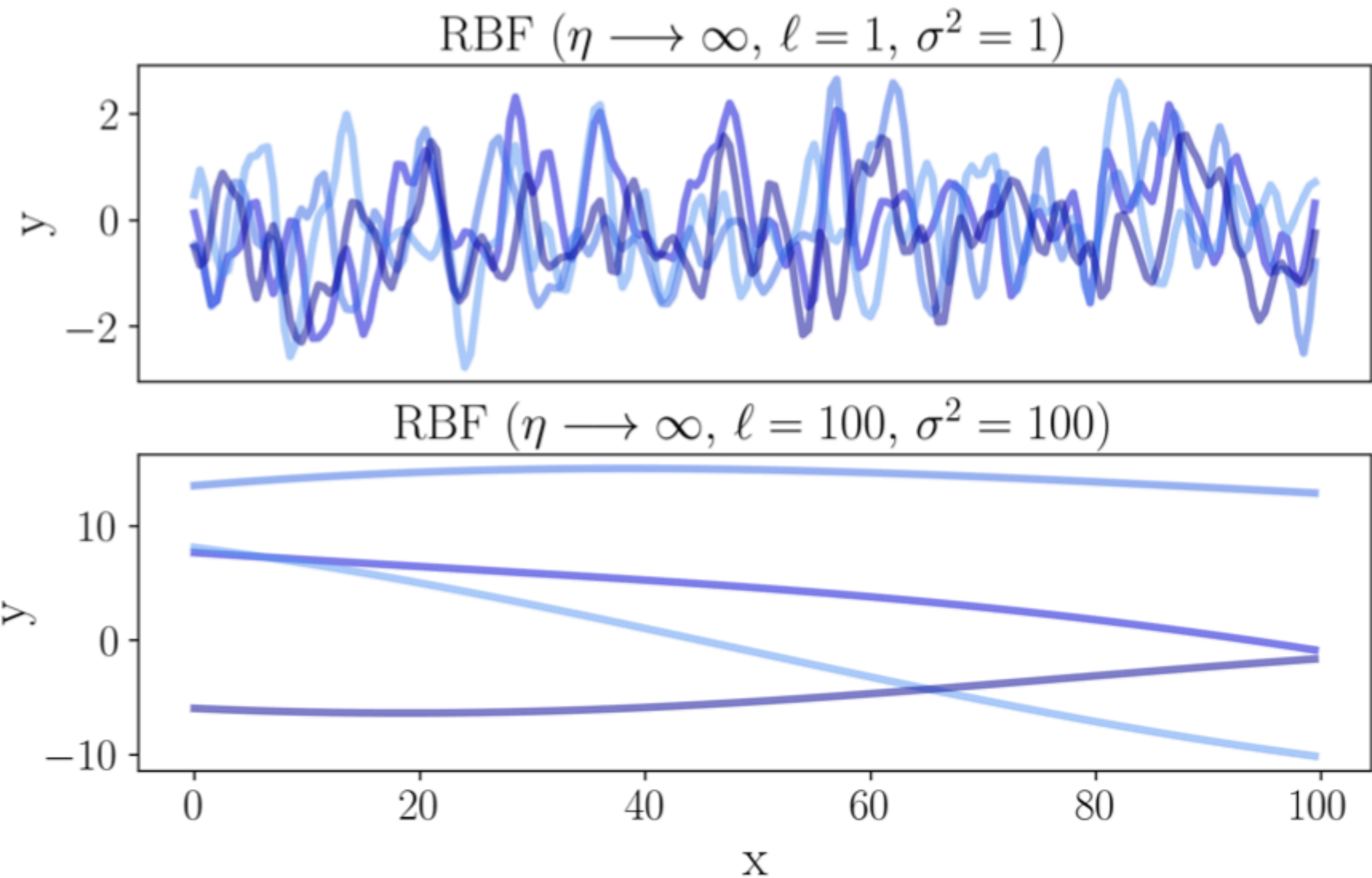
$l$ :Length scale(topical scale of correlations in the data across frequency)

$\eta$ :spectral parameter(It determines the overall "smoothness" of the data)

# $\sigma^2$ and $l$ of Matern kernel

- Larger  $\sigma^2$  (Variance)  
→ The signal is stronger.
- Larger  $l$  (coherence scale)  
→ The signal is more correlated in frequency

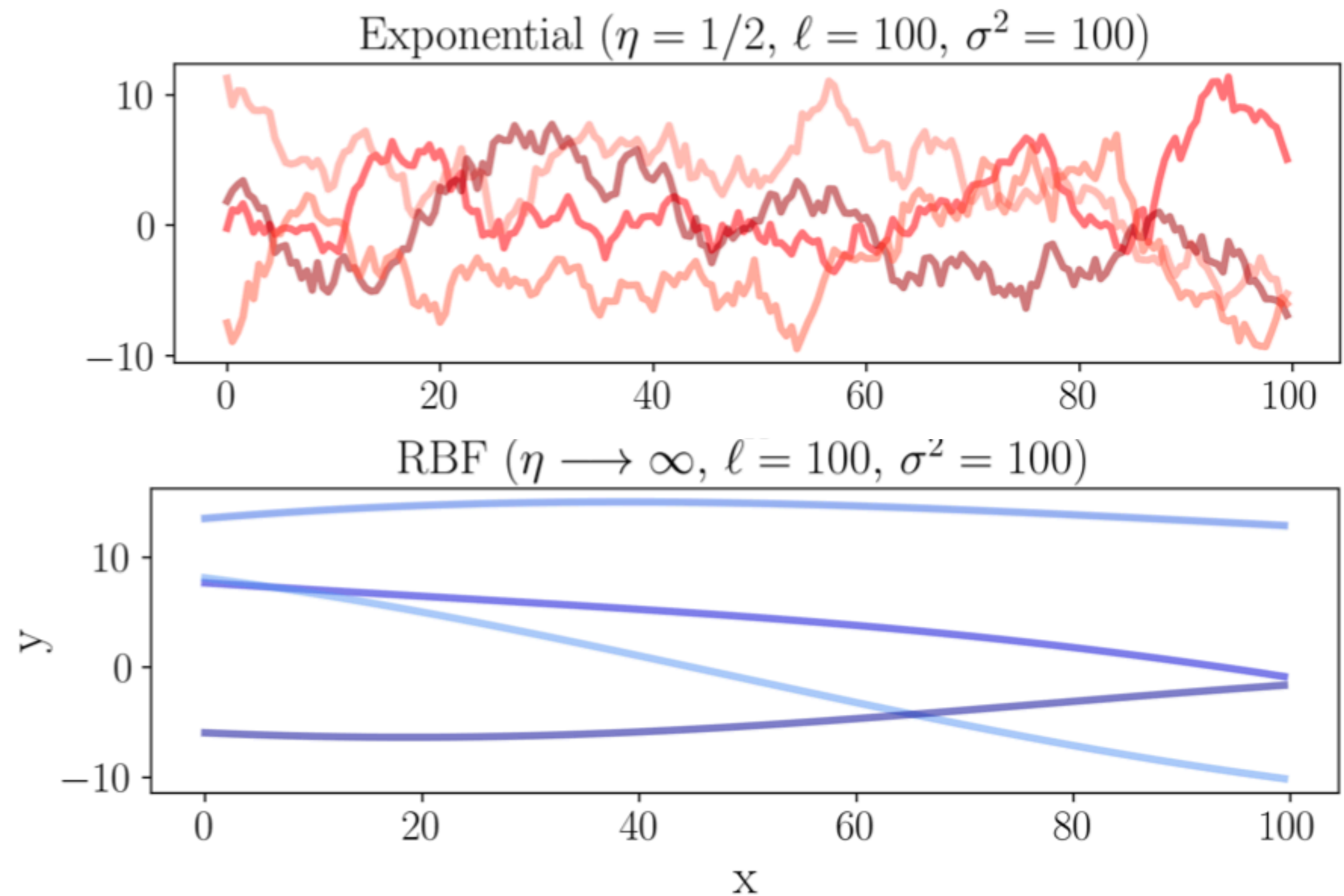
randomly generated data plots with Matern kernels,  
with shown parameters



# $\eta$ of Matern kernel

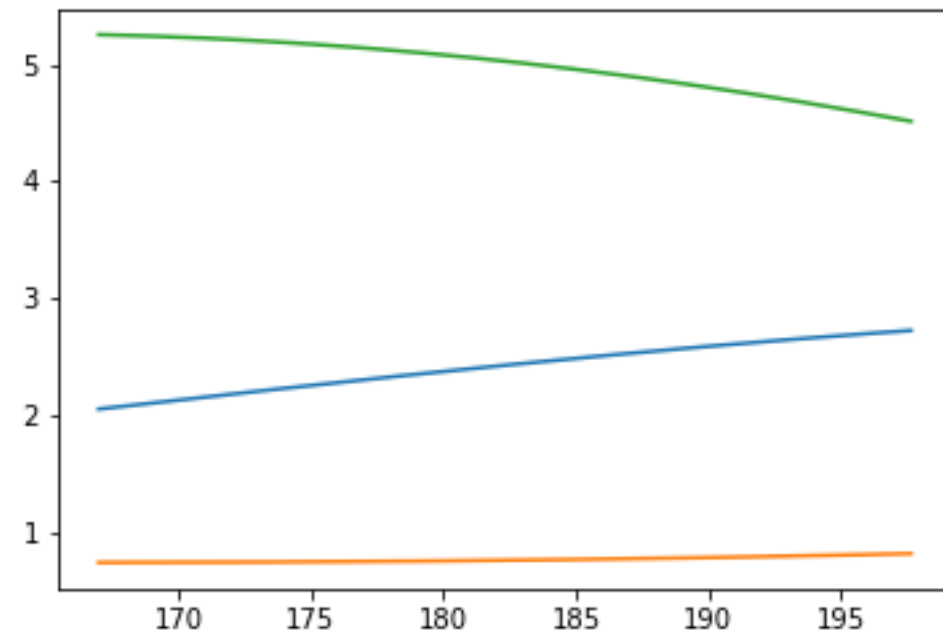
- Larger  $\eta$   
→ data is spectrally smoother
- FG is spectrally smoother!

randomly generated data plots with Matern kernels,  
with shown parameters

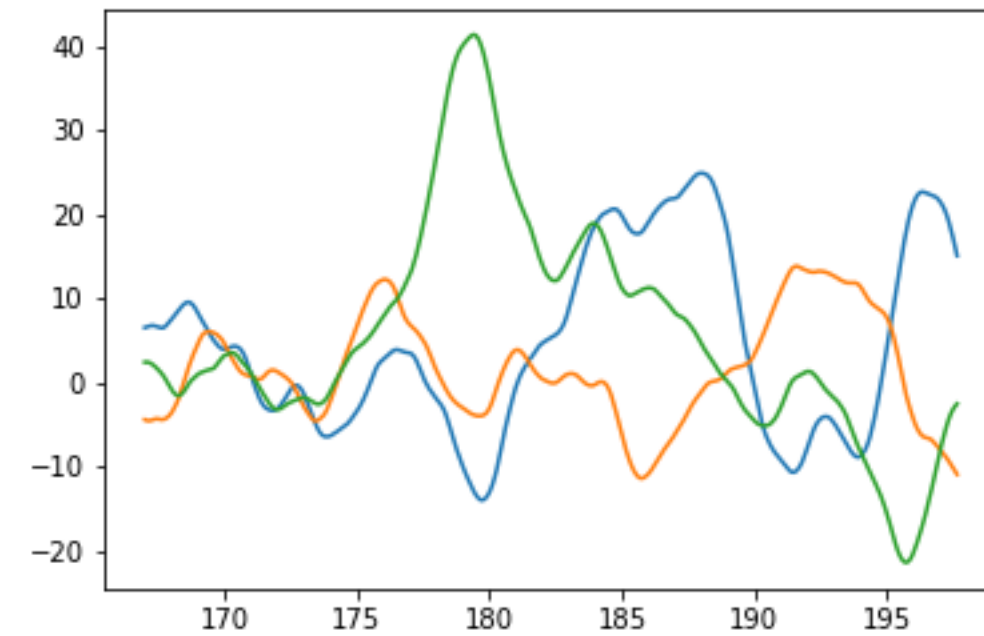


# Components

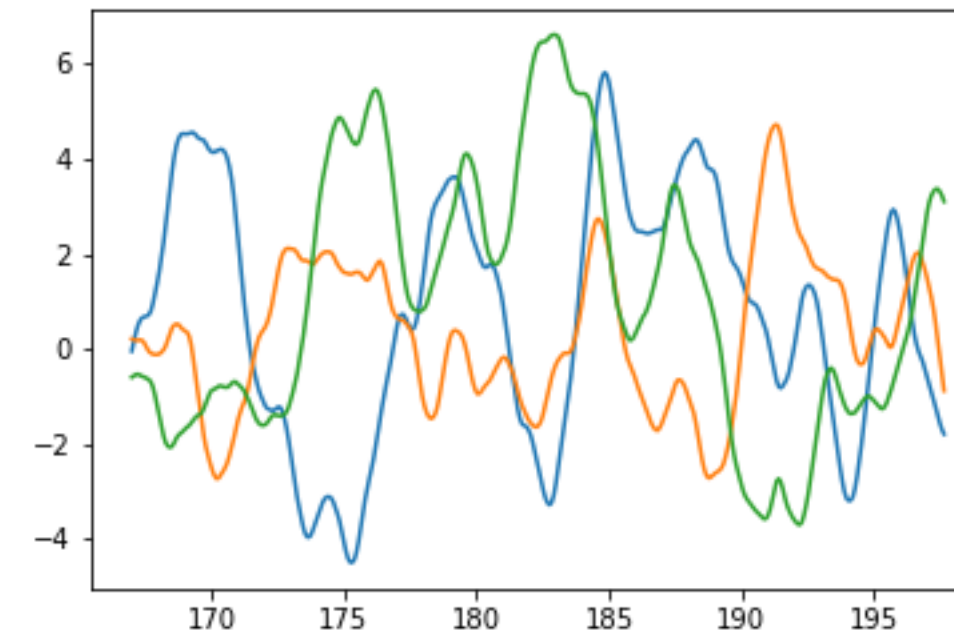
Smooth FG



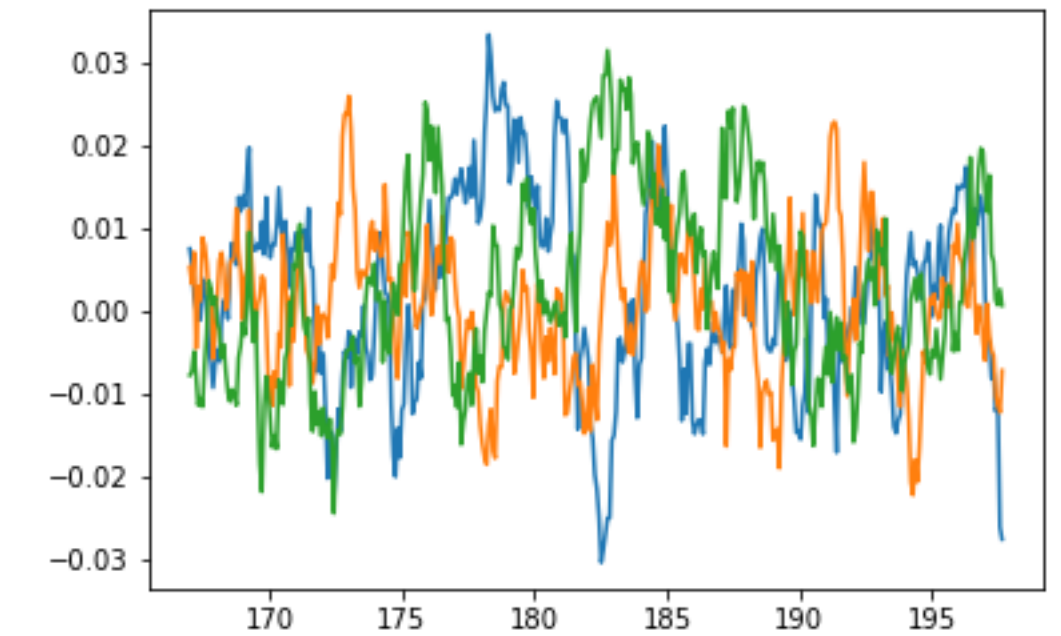
non-smooth FG



Additional component

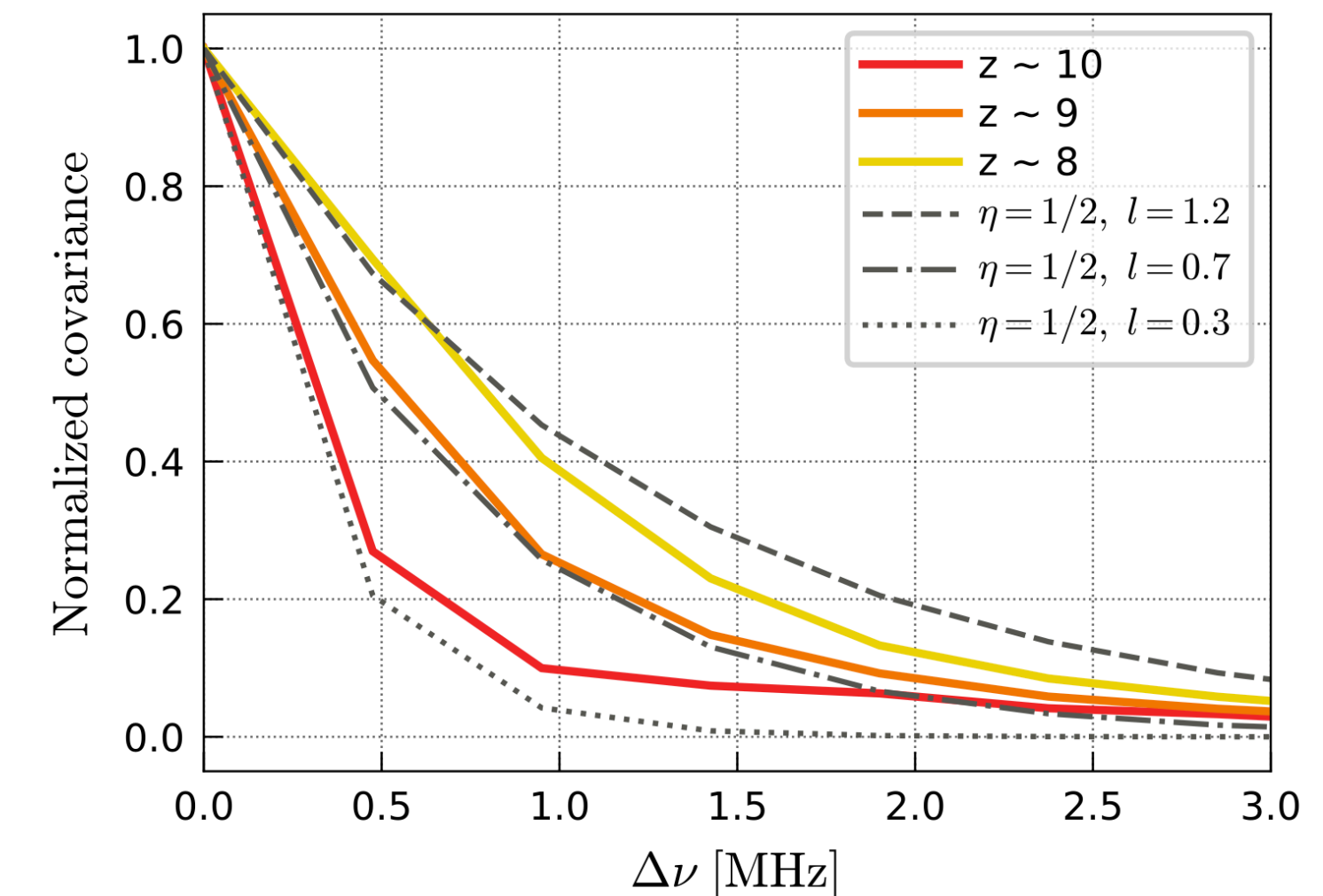


HI



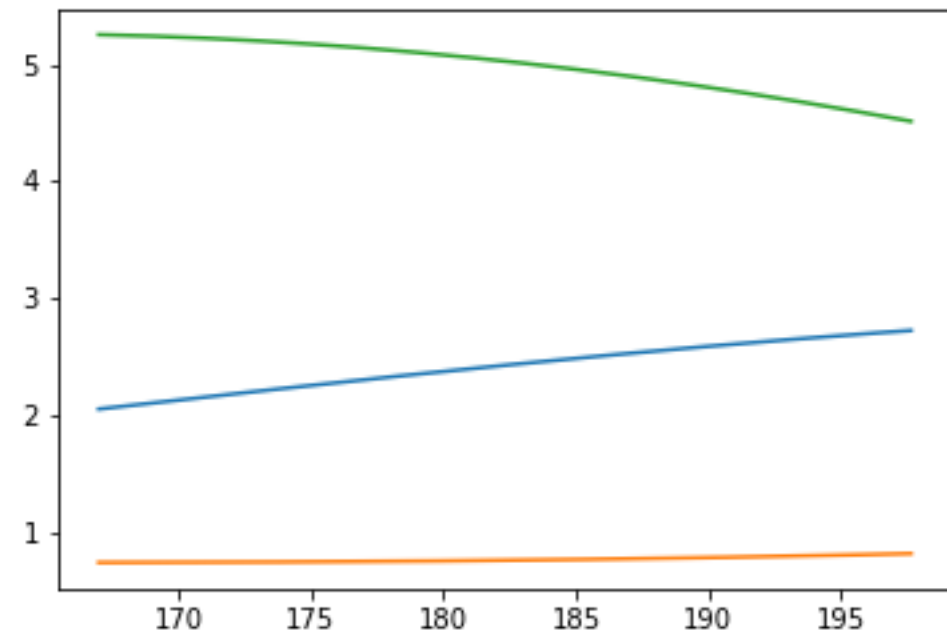
Mertens et al 2018

- Covariance of 21cm line  
→ Well approximated by  $\eta = 1/2$
- We don't know what are the best FG kernels for MWA!  
→ Compare some kernels sets using Bayesian evidence



# Applied kernel and its parameter priors

Smooth FG

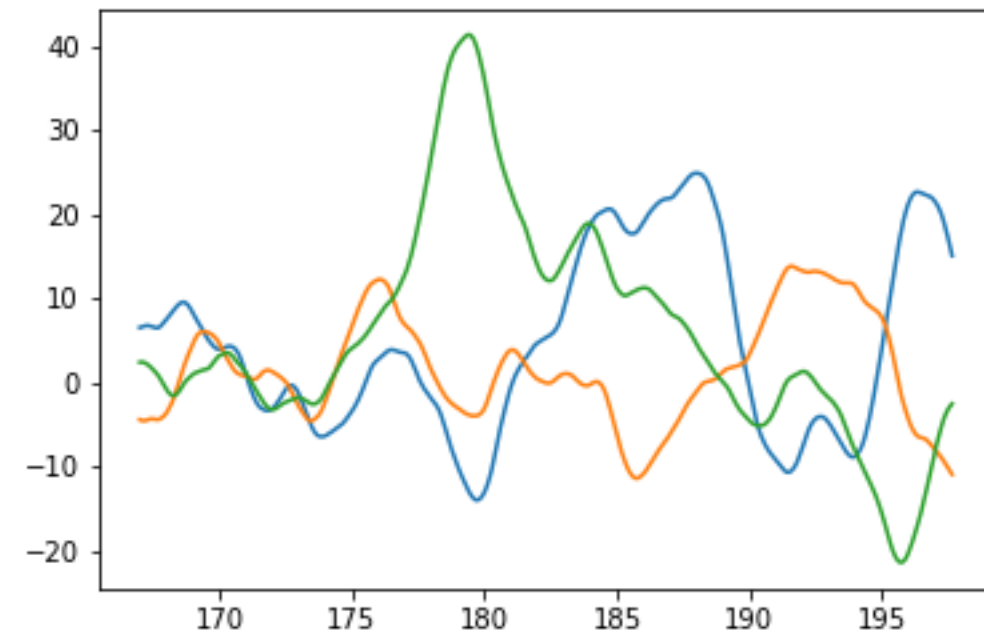


$$\eta \rightarrow \infty$$

$$10 < l < 100[\text{MHz}]$$

—

non-smooth FG

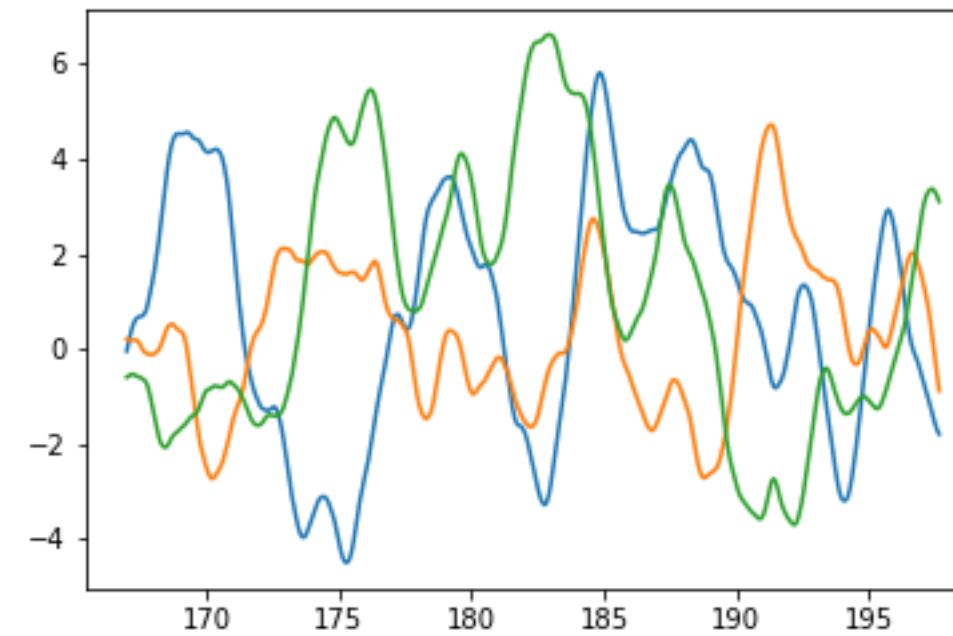


$$\eta = 5/2 \text{ or } 3/2$$

$$1 < l < 10[\text{MHz}]$$

—

Additional component

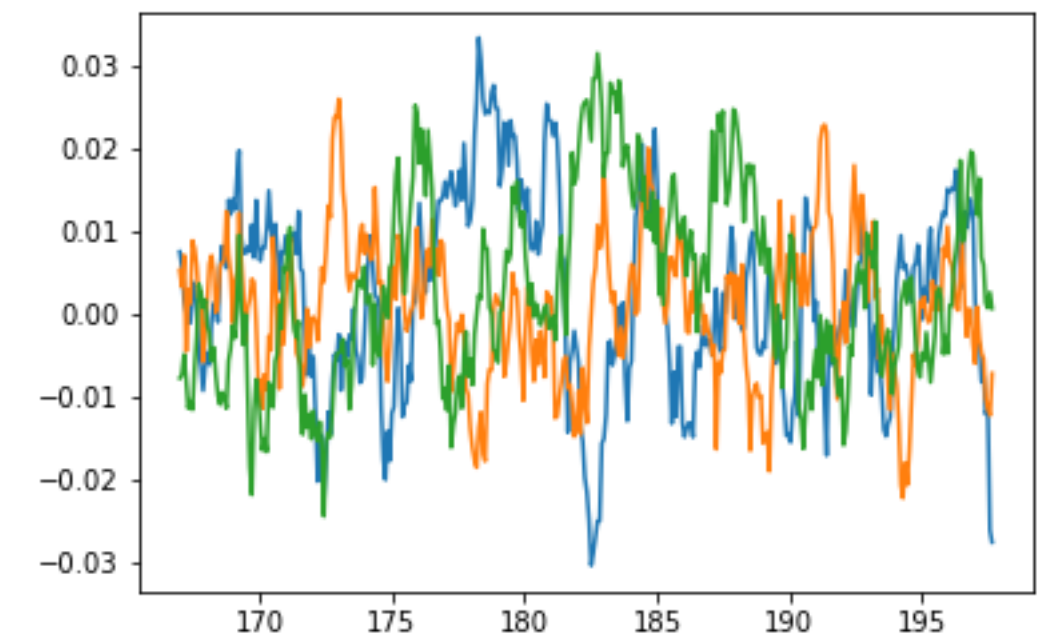


$$\eta = 5/2 \text{ or } 3/2$$

$$0.64 < l < 1.92[\text{MHz}]$$

—

HI



$$\eta = 1/2$$

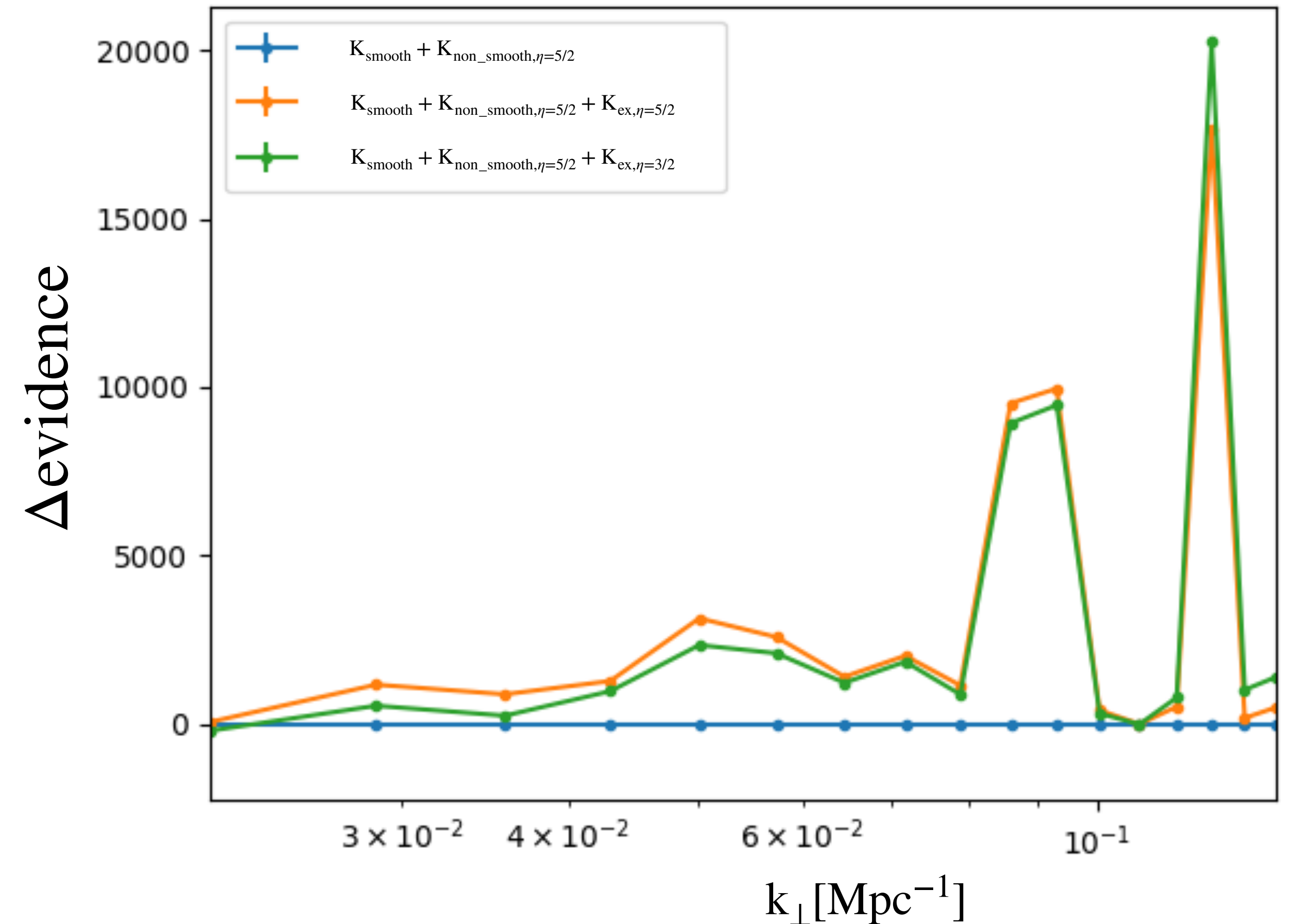
$$0.1 < l < 1.2[\text{MHz}]$$

$$10^{-6} < \sigma^2 < 0.5[\text{Jy}^2]$$

- Apply GPR to the data with each kernel set, and compare its evidence  
-MWA Observational data (high band observation in 2015 (2h@EoR0))

# Result 1 ~Best kernel for MWA~

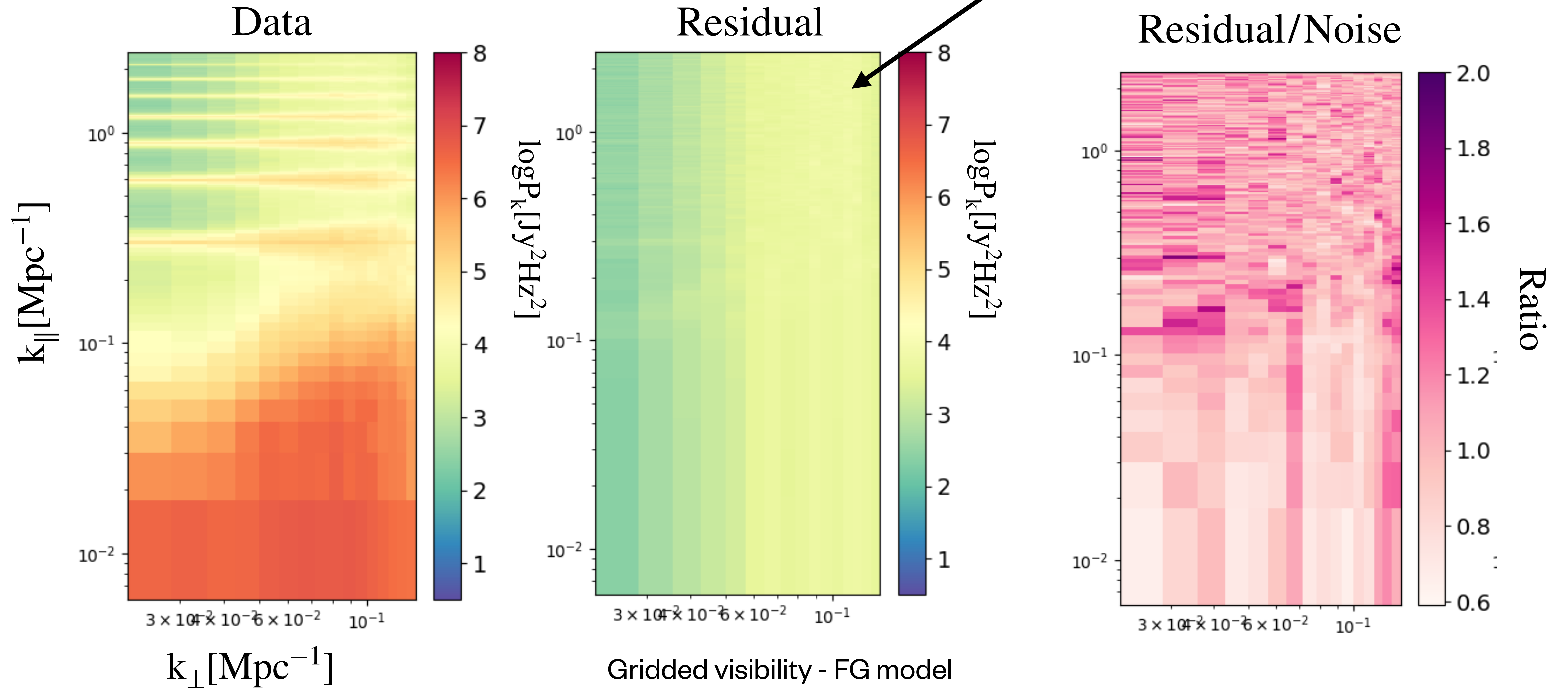
- Compare Bayesian Evidence of different kernel sets
- $\Delta\text{evidence} = \log Z_{\text{model}} - \log Z_{\text{fiducial}}$ 
  - $|\Delta\text{evidence}| > 10$ :  
→ Strong difference
  - Higher  $\Delta$  evidence  
→ Better model
- Chosen kernels are depend on the value of  $K_{\perp}$





# Result 2~2D power spectrum~

Coarse band harmonics disappeared



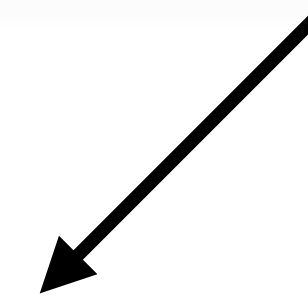
# Signal loss

Additional component

$0.64 < l < 1.92$ [MHz]

HI

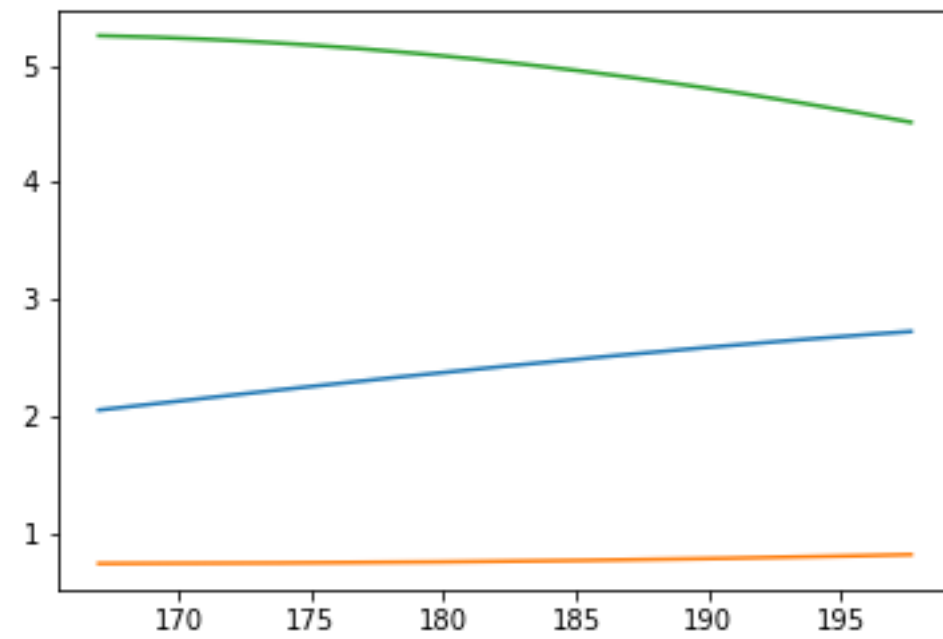
$0.1 < l < 1.2$ [MHz]



- Additional component has similar coherence scale to HI kernel  
→ Subtracting additional component may cause signal loss
- Generate Multivariate Gaussian Distribution with K as signal component  
→ Apply GPR based FG removal method to generated data and recover HI power spectrum

# Results 3 ~signal loss~

$K_{\text{smooth}}$

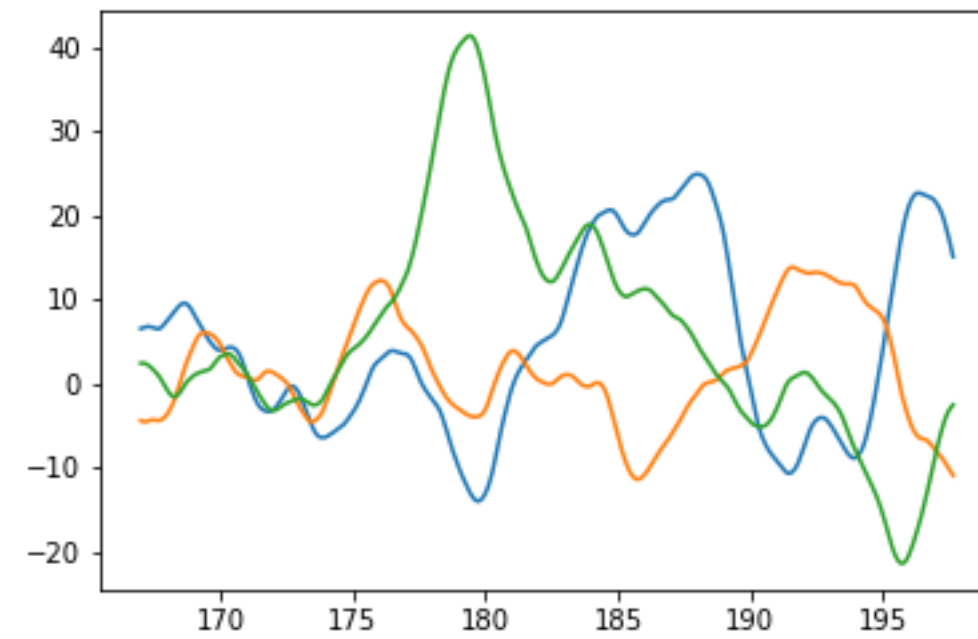


$$\eta \rightarrow \infty$$

$$l = 100[\text{MHz}]$$

$$\sigma^2 = 5[\text{Jy}^2]$$

$K_{\text{non\_smooth}, \eta=5/2}$

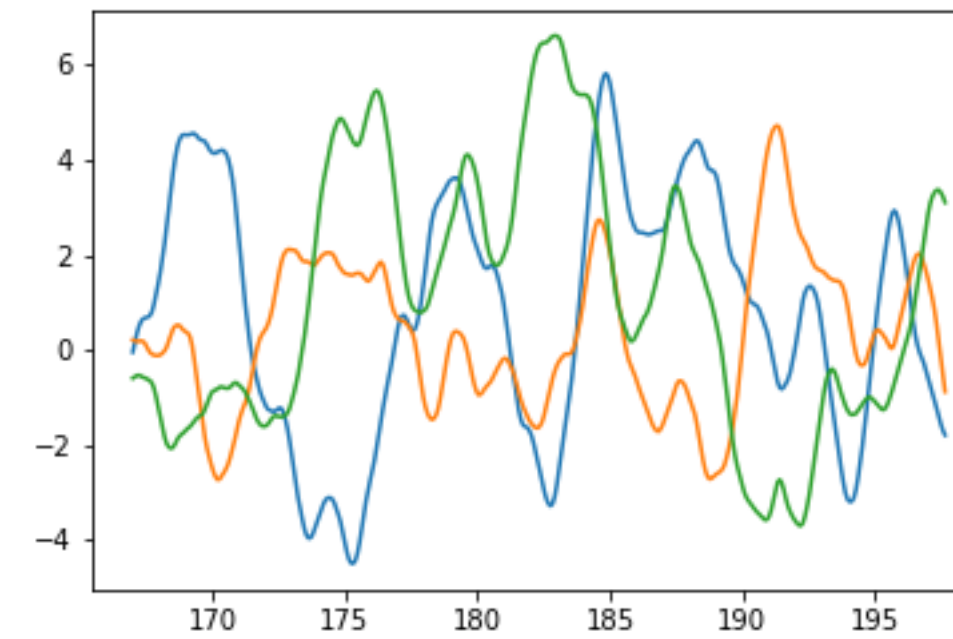


$$\eta = 5/2$$

$$l = 5[\text{MHz}]$$

$$\sigma^2 = 100[\text{Jy}^2]$$

$K_{\text{add}, \eta=5/2}$

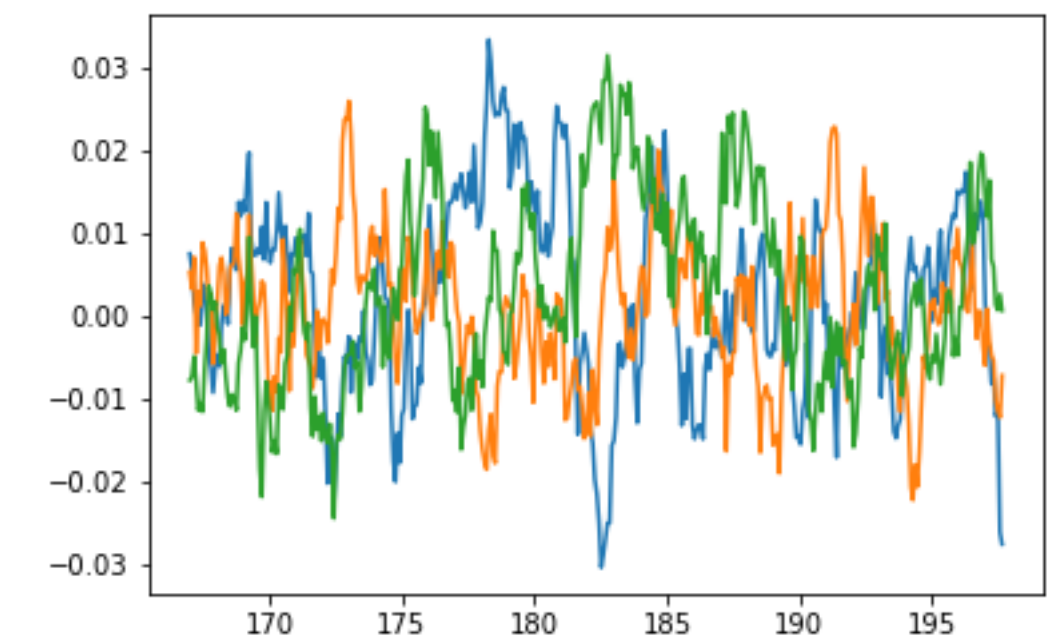


$$\eta = 5/2$$

$$l = 1.28[\text{MHz}]$$

$$\sigma^2 = 5[\text{Jy}^2]$$

$K_{\text{HI}}$



$$\eta = 1/2$$

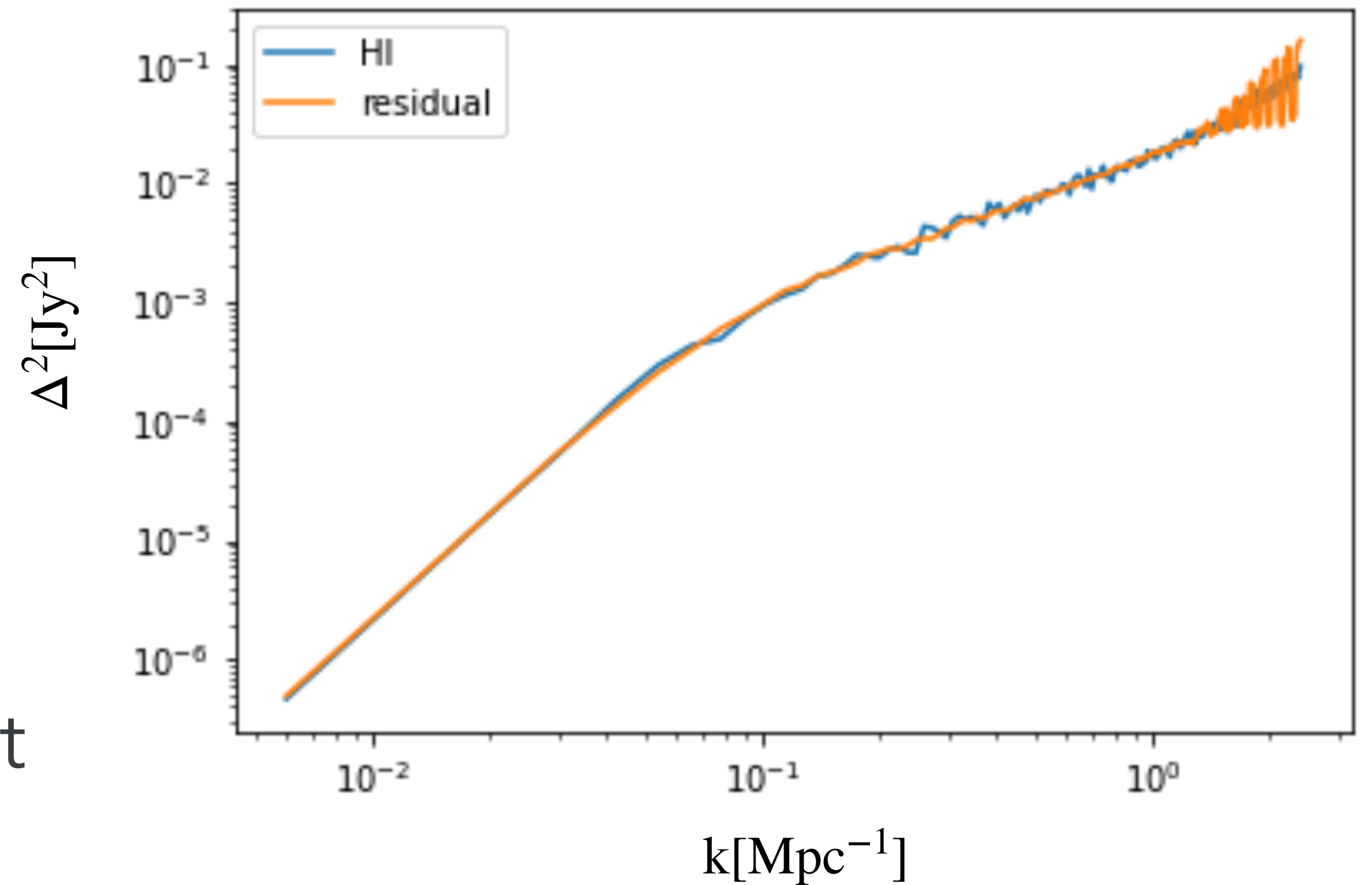
$$l = 0.75[\text{MHz}]$$

$$\sigma^2 = 0.0001[\text{Jy}^2]$$

- Combine these signals together as simulated signal  
→ GPR based FG removal

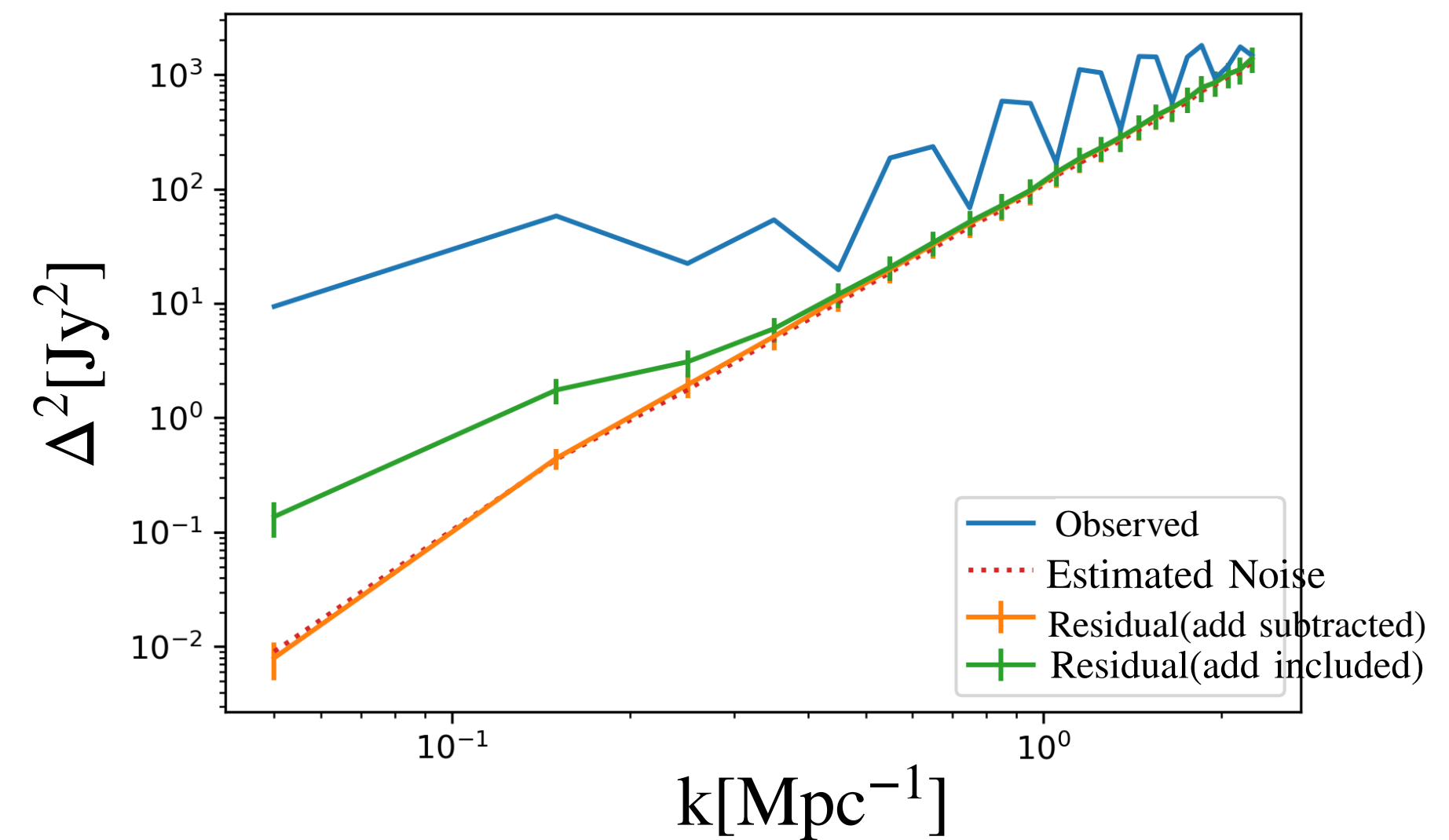
# Results 3~Signal loss~

- HI power spectrum VS Residual
- If GPR work well
  - Ideal parameters are chosen
  - Signals are completely represented by kernel
  - No signal loss even if we subtract additional component



# Conclusion

- I applied GPR based FG removal to MWA observational data(2h in 2015)
  - Kernel set with Additional power kernel has better Bayesian evidence
  - Residual is smaller than observed data
    - $\sim 10^2$  [Jy<sup>2</sup>] (additional component included)
    - $\sim 10^3$  [Jy<sup>2</sup>] (additional component subtracted)



- Maybe no signal loss even if we subtract additional component

